

# Summary beam dynamics group

Incomplete review of material and response on initial questionnaire

## Halo definition

- part of beam exceeding  $\sim 3 \sigma$  that needs to be scraped
- tails exceeding Gaussian tails
- halo is whatever bothers you!

# How complete is our understanding of halo mechanisms?

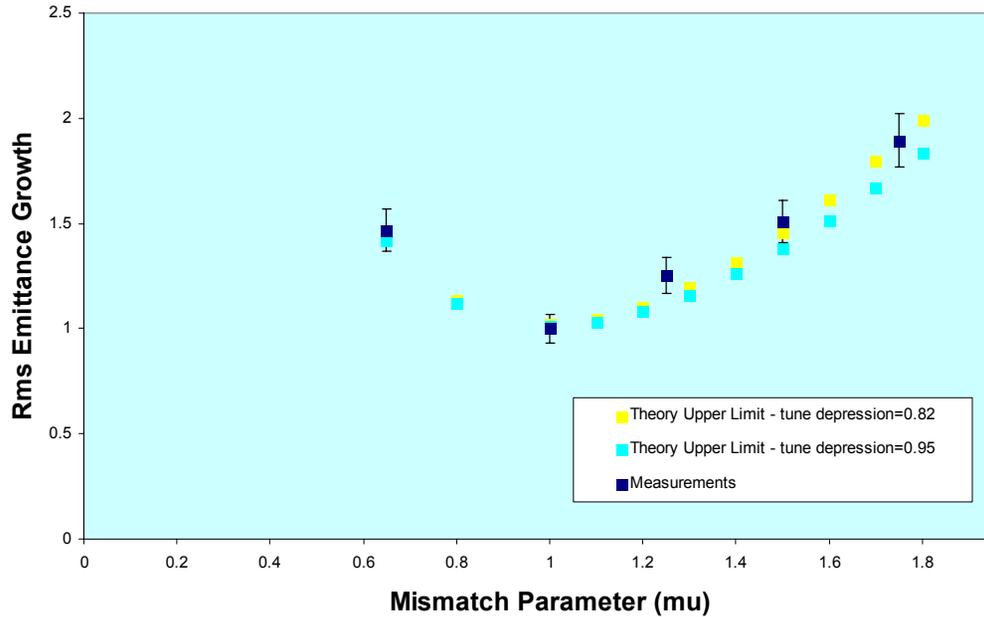
## “Resonant halo”

### Coherent motion drives halo

1. Envelope mismatch (parametric 2:1 halo) R + L  
(T.Wangler, N.Pichoff, F.Gerigk, J.Qiang,
2. Random lattice errors (envelope or rotation) L  
(F. Gerigk, P.Ostroumov, R.Kishek)
3. Dipole motion mismatch \*) damped by halo R  
or transverse (dipole) instability  
(V.Danilov) – *to be explored further for SNS ring*
4. Anisotropy resonance/instability (2:2) ~ low halo R + L  
(JM.Lagniel, J.Qiang, D.Jeon)

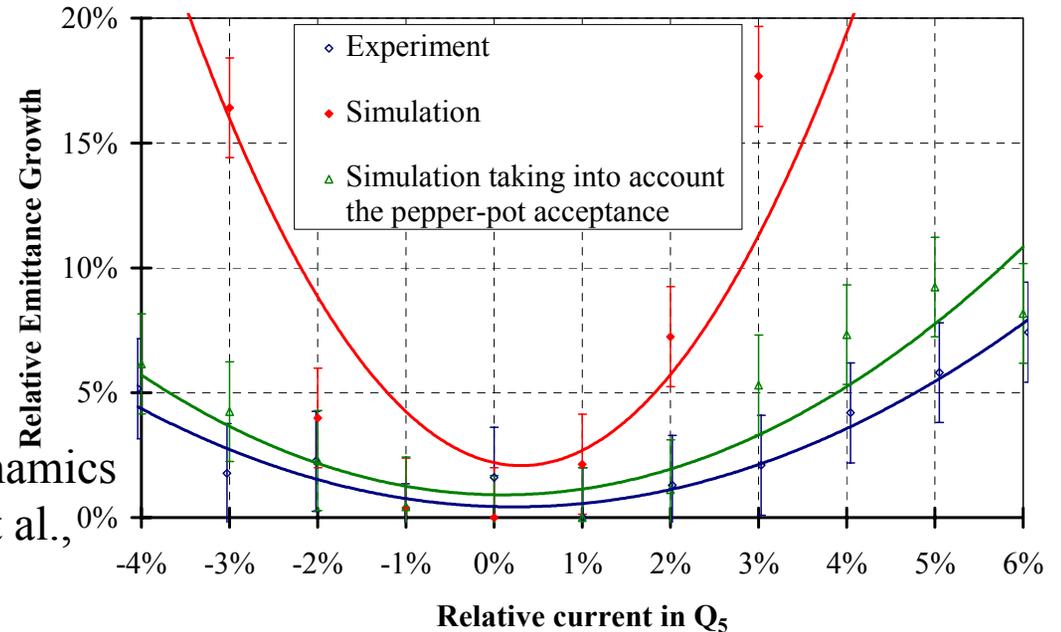
\*) *off-centered: in beam-beam group synchrotron radiation damps for electrons – not for hadrons*

RMS EMITTANCE GROWTH AT SCANNER #45 - 75 mA - BREATHING MODE



LEDA exp.  
T. Wangler, J.Qiang et al.

Halo growth agrees with experiment at the level of rms emittance growth  
– *halo size open*



CEA: Measurement of space charge-dynamics effects in a FODO channel, N. Pichoff et al., EPAC98

Have overlooked early – 1985 ... - experiment  
(M. Tiefenbach et al.) + simulation (O. Andersen) at LBL  
(heavy ion fusion program)

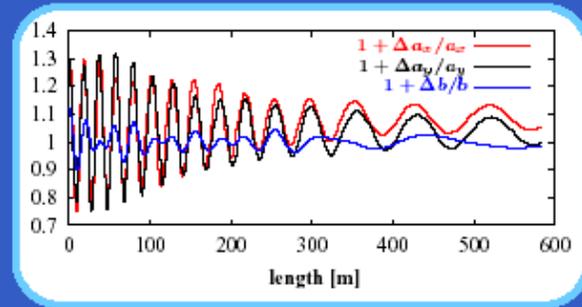
- in a ~ 50 quad transport channel
- to demonstrate structure resonances (envelope instability  $\sigma_0 > 90^\circ$ , etc.) and compare with theory (1978-82)
- but measured also halo by mismatch for strong space charge

Halo growth by random error in worst cases comparable with initial mismatch

## Transverse Errors (radial deviation)

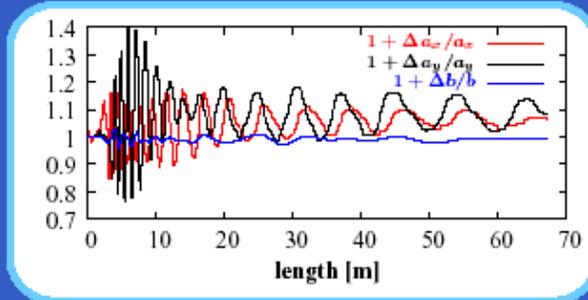
F. Gerigk

**Halo in a nutshell:** Halo particles are generated by parametric resonances between single particles and the oscillations of a mismatched beam core.



**30 % fast mode excitation in a high-intensity linac**

### Radial deviation for the full Linac4:



**$\Delta a_x$ ,  $\Delta a_y$ , and  $\Delta b$  for the worst case in y (1% gradient variation)**

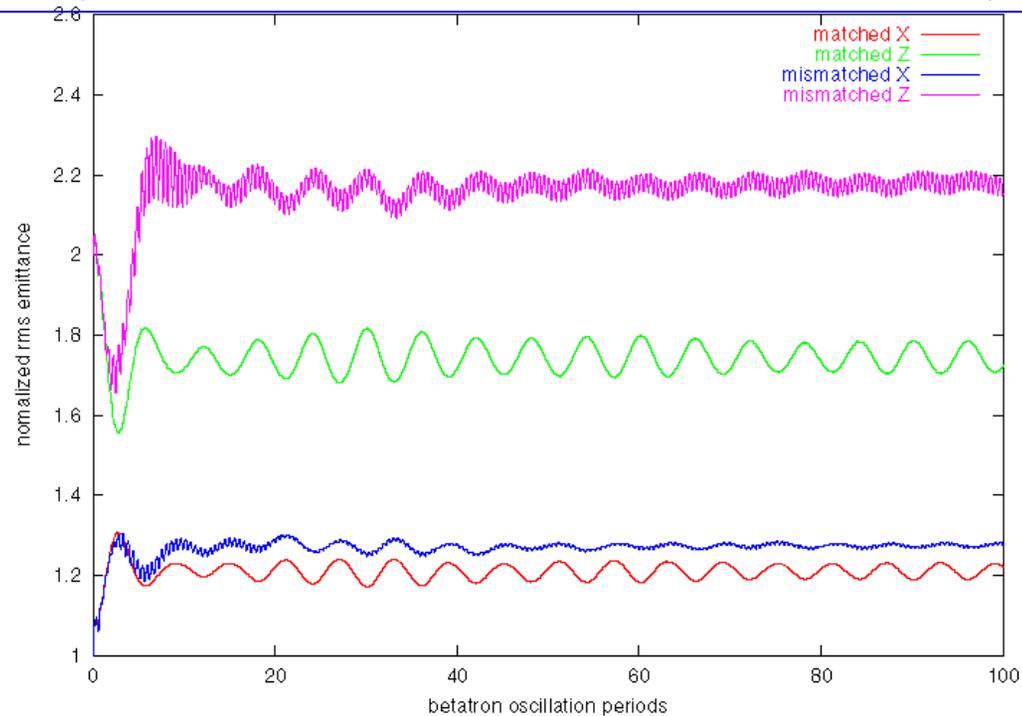
- Statistical errors also result in core oscillation!
- The highest radial deviation compared to the matched case is observed after the chopper line in the first DTL tank!



Interplay of non-equipartitioned and envelope mismatch may drive beam still further away from equipartition

### RMS Emittance Evolution in a Matched and Mismatched Anisotropic Beam

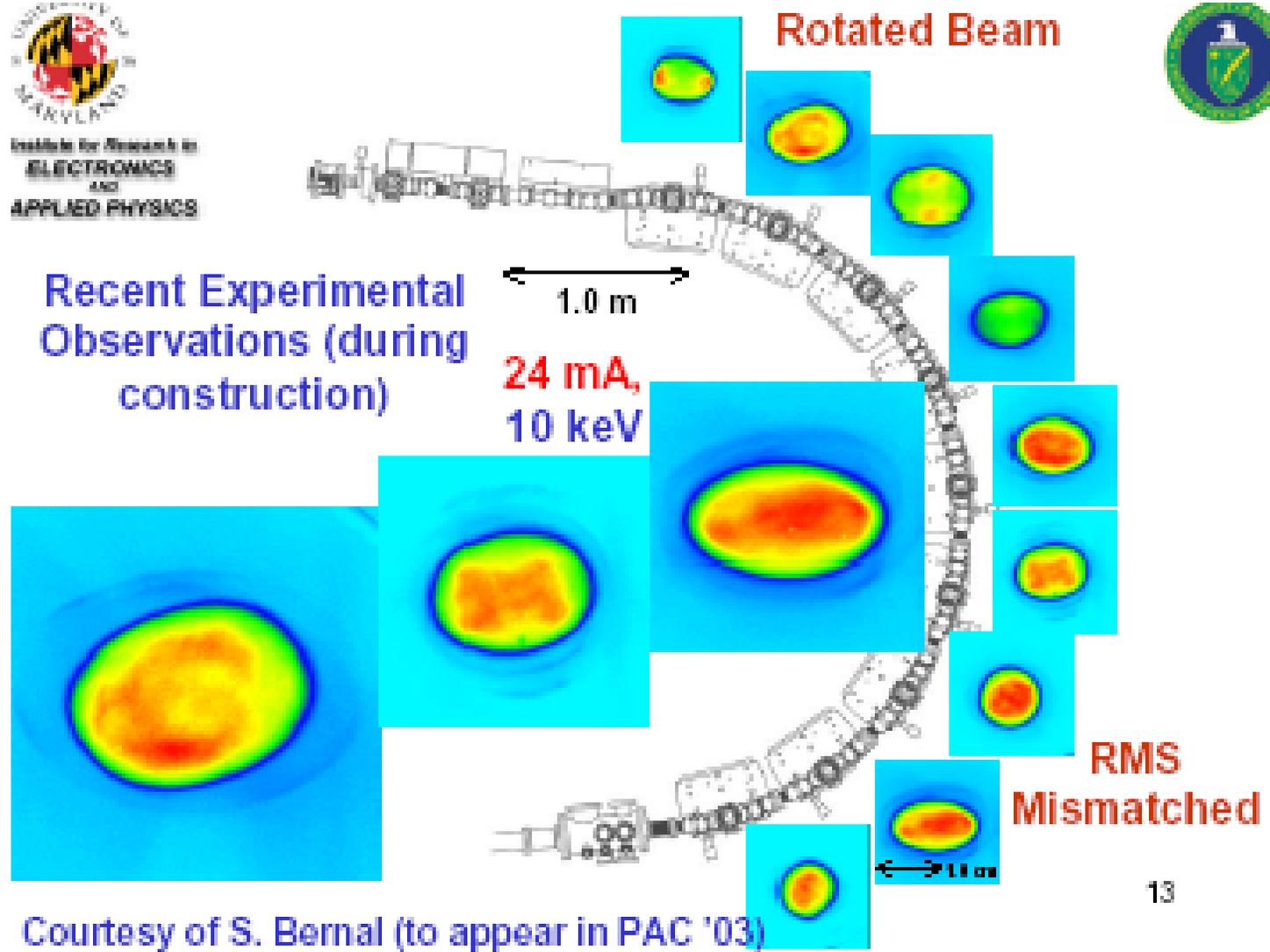
( $K_z0/K_x0 = 1.025$ ,  $K_x/K_x0=0.6$ ,  $K_z/K_x = 1.26$ ,  $E_z/E_x = 2$ , Gaussian Distribution)



*J. Qiang et al.*



Institute for Research in  
ELECTROMAGNETICS  
AND  
APPLIED PHYSICS



Recent Experimental  
Observations (during  
construction)

24 mA,  
10 keV

Rotated Beam

RMS  
Mismatched

Courtesy of S. Bernal (to appear in PAC '03)

# How complete is our understanding of halo mechanism?

## “Resonant halo” cont’d

External driving term (*lattice harmonic*) is source

### 1. Second order resonance

(A.Fedotov, SNS-ring)

< 1ms

R

→ pronounced coherent effect / advantageous

### 2. Higher order resonance

R

a) “space charge octupole”  
(S.Igarashi, S.Machida)

< 1 ms

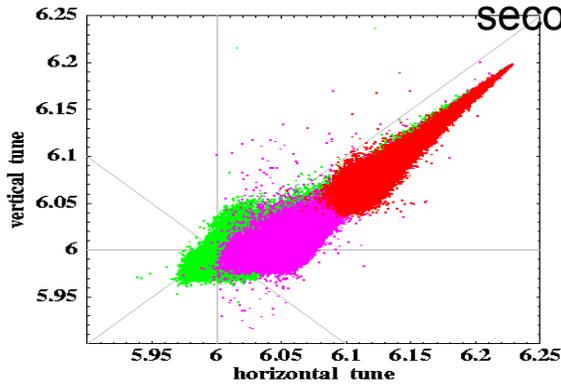
b) nonlinearity + synchrotron motion  
(G.Franchetti, I.Hofmann)

> 10 ... 1000 ms

→ ~ negligible coherent effect for Gaussian beams  
allows analytical space charge simulation

c) tune modulation – different kinds

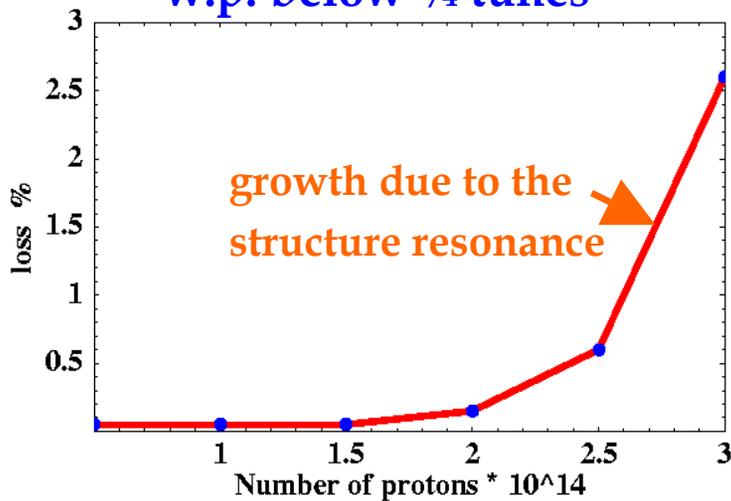
second order resonance



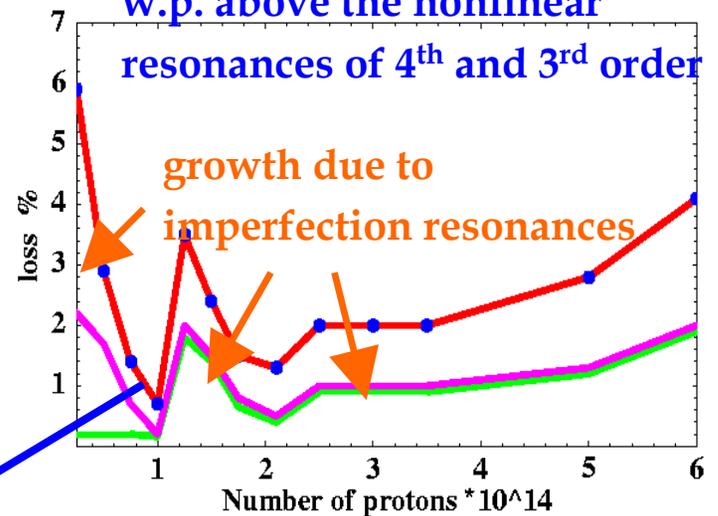
## SNS examples

realistic simulations with the UAL code

w.p. below  $\frac{1}{4}$  tunes



w.p. above the nonlinear resonances of 4<sup>th</sup> and 3<sup>rd</sup> order

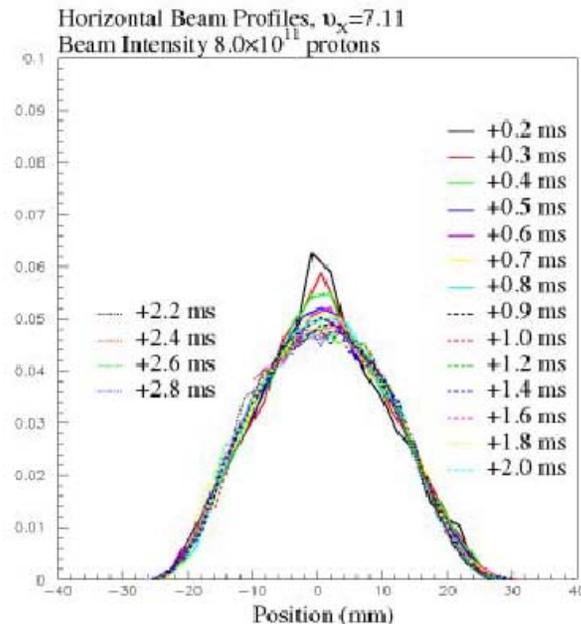
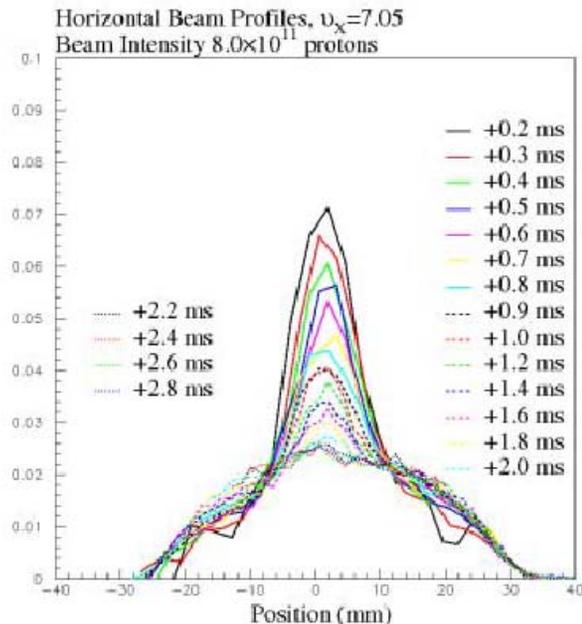


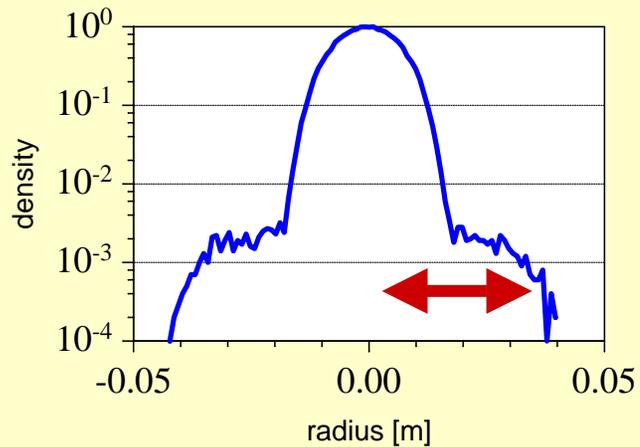
requires resonances correction

## KEK main Synchrotron: observed and simulated shoulders (octupolar resonance)

## Horizontal Beam Profiles

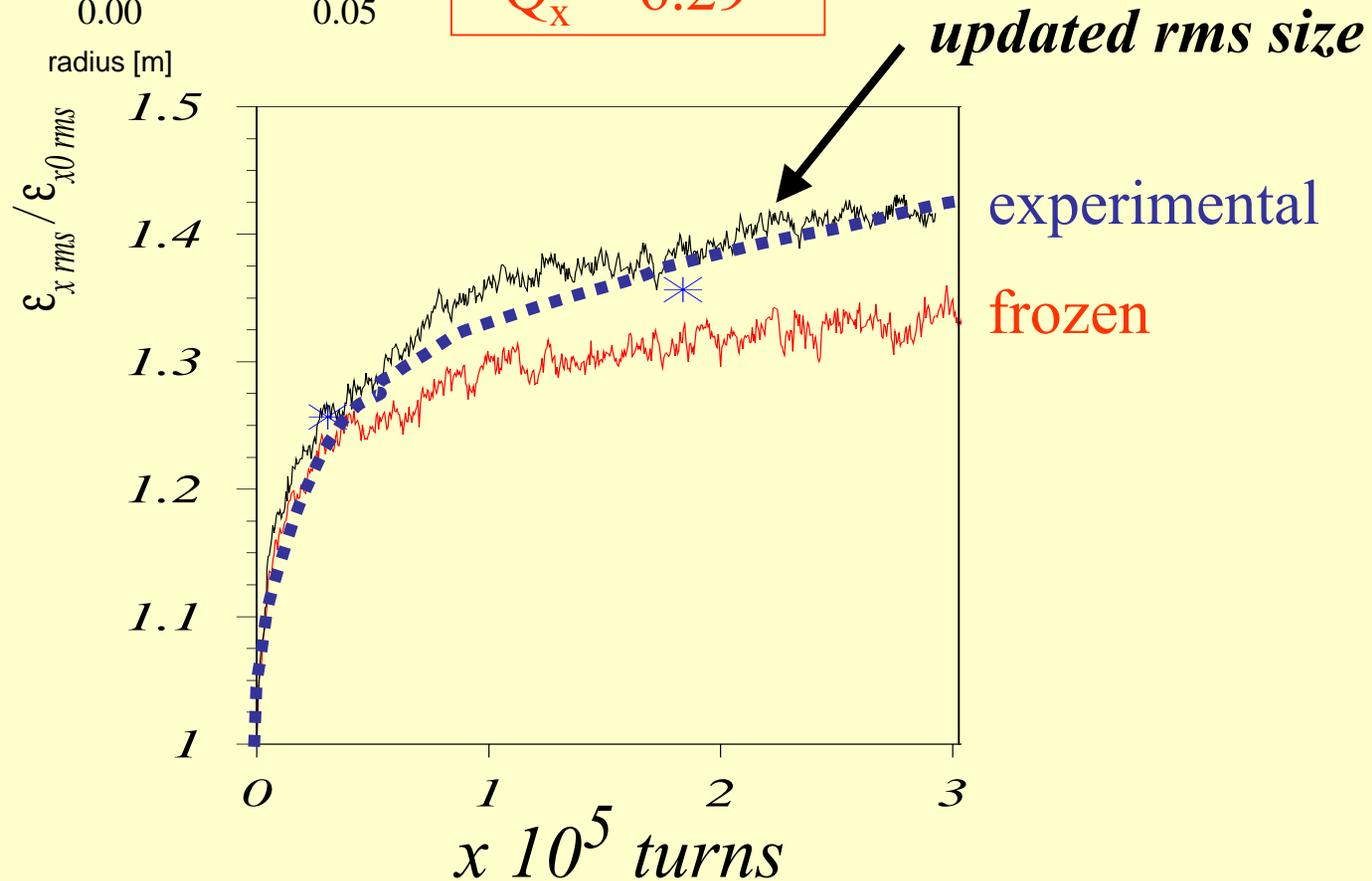
0.2 ms ~ 2.8 ms after Injection

Beam Intensity  $8.0 \times 10^{11}$  protons



*CERN-PS halo benchmarking experiment:*  
halo unbounded and particle amplitudes go  
in and out over many synchrotron periods

$$Q_x = 6.29$$



# How complete is our understanding of halo mechanism?

## Scattering induced halos (*non-resonant*)

(N.Pichoff, G.Turchetti)

R + L

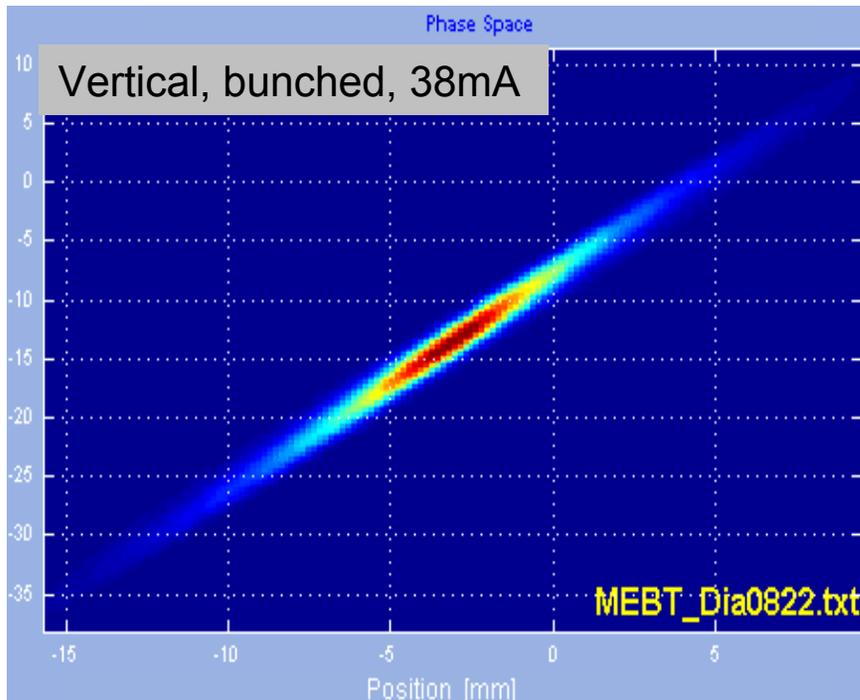
*IBS only in R*

## Further machine specific sources (*non-resonant*)

1. RF bunching mismatch responsible for ~ 800% emittance growth in LANSCE
2. In SNS and J-PARC MEBT line “spherical aberration” (octupolar force from space charge or RF)  
observed in experiment and modeled by simulation (D.Jeon, M.Ikegami)
3. H- stripping / Interaction with residual gas / Image effects etc.

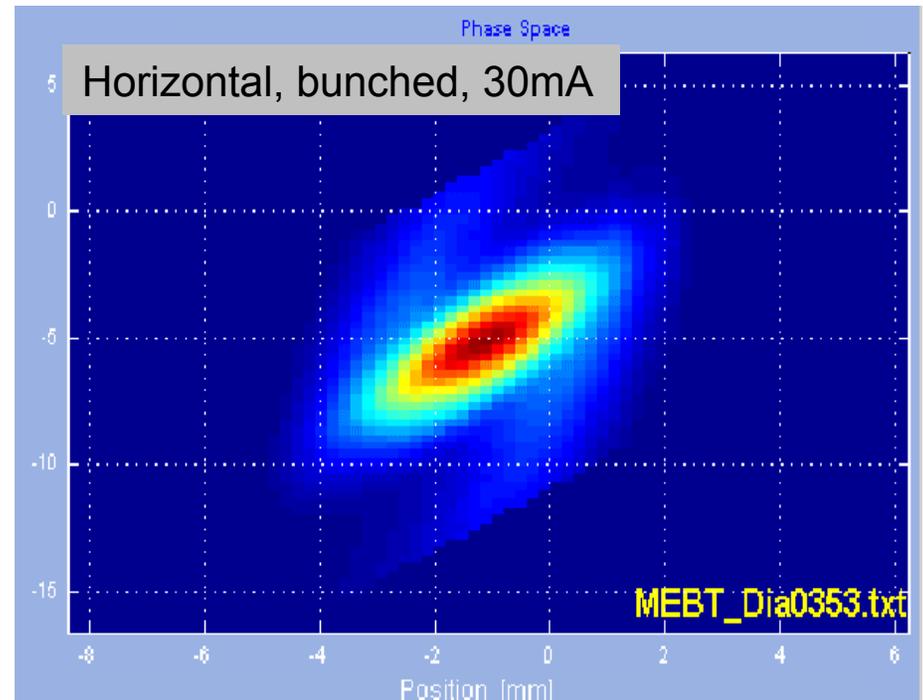
# Transverse emittance at the MEBT exit (II)

- Transverse emittance for optimized settings is within specified limits ( $<0.3 \text{ } \mu\text{m mrad}$ )



normalized r.m.s. =  $0.3 \text{ } \mu\text{m mrad}$

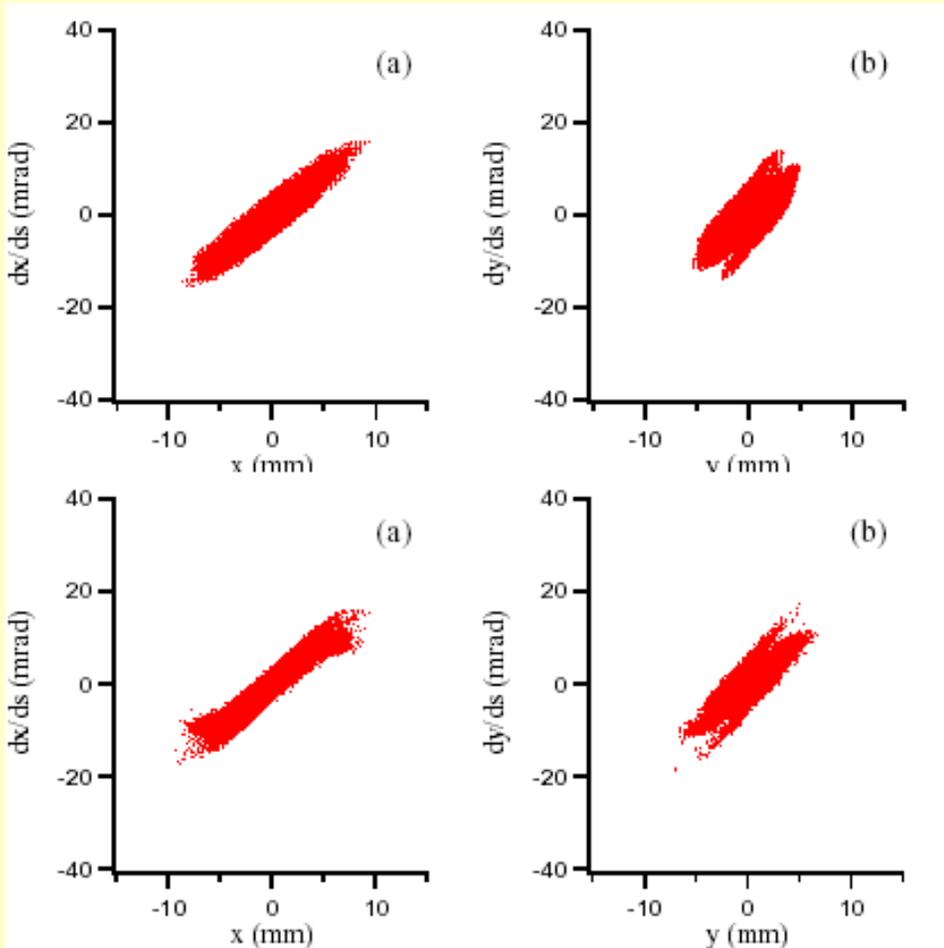
normalized enclosed 90% =  $1.8 \text{ } \mu\text{m mrad}$



normalized r.m.s. =  $0.24 \text{ } \mu\text{m mrad}$

normalized enclosed 90% =  $1.58 \text{ } \mu\text{m mrad}$

# Comparison of phase-space profile (GA)



Experiment

**J-PARC Linac**  
M. Ikegami

Simulation

# How complete is our understanding of halo mechanism?

## Studies on specific machines: halo impact

1. BNL Linac/AGS 1MW upgrade for neutrino factory discussing effect of stripping foils and many others (W. Weng)
2. RIA Linac (P. Ostroumov) – **basically halo free design!**
3. CERN sc Linac (F. Gerigk)
4. J-PARC Linac (M. Ikegami)

## What determines size of halo?

- Parametric 2:1 halo:  $10^{-4}$  may reach  $\sim 3 \sigma \rightarrow 5-7 \sigma$   
(MM=1.3, upper limit?)
- Coupling long-transv may enhance halo size
- Random error halo:  $\sim$  same ?  
*more work may be needed*
- Space charge octupole (KEK) shoulders
- Nonlinearity + synchrotron oscillation (CERN-PS)  $3 \sigma \rightarrow$  unbounded

## Dominant mechanisms of halo production

Linac:

1. Initial mismatch from poor matching between sections – in principle cured by sufficient diagnostics
2. Random errors (gradients, misalignment, ...) – do we find pronounced parametric halo?
  - a. profiles of simulation for some error sets of CERN-SPL indicate halo up to  $5\sigma$ ?
  - b. need more systematic work!
3. Source halo, scattering, ...?

Ring:

1. Initial mismatch “parametric halo” with bunch transfer should be possible to minimize
2. Nonlinear resonances – depends on time scale 1 ms - 100-1000 ms

# Conclusion

Lots of work left for

- experiments +
- design-close dynamics studies