

# Stochastic Cooling for RHIC

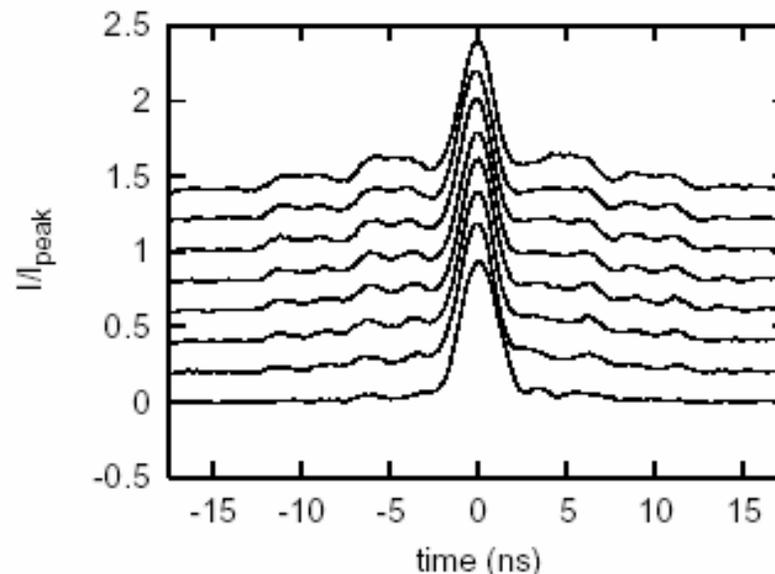
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- Introduction
- Results from last year
- Coherent line problem
- Hardware for this year
- Plan for this run
- Future ideas

# Introduction

We plan to implement “Halo Cooling” to counteract IBS driven debunching.  
Improve integrated luminosity by a significant factor within the next few years.

Debunching reduces useful luminosity by lengthening vertex.



# Bunched-beam Stochastic Cooling

- What would be required,
  - Cooling time would have to be commensurate with de-bunching time, ~ few hours
  - Cool only large  $\Delta P$  particles (halo cooling)
- Consider coasting beam theory (full bucket)

$$\frac{1}{\tau} = \frac{W}{N_{\text{eff}}} \left[ 2g(1 - \tilde{M}^{-2}) - g^2(M + U) \right] \quad g_{\text{optimum}} = \frac{1 - \tilde{M}^{-2}}{M + U} \cong \frac{1}{5}$$

$$\frac{1}{\tau_{\text{opt}}} = \frac{W}{N_{\text{eff}}} \left[ \frac{(1 - \tilde{M}^{-2})^2}{M + U} \right] = \frac{1}{3000 \text{ sec}} \quad N_{\text{effective}} = \frac{10^9}{1.5\text{m}} 3830\text{m} = 2.5 \times 10^{12}$$

- Why wasn't stochastic cooling in the base line design for RHIC?
- High frequency **bunched-beam** stochastic cooling is required

# Highlights from last year

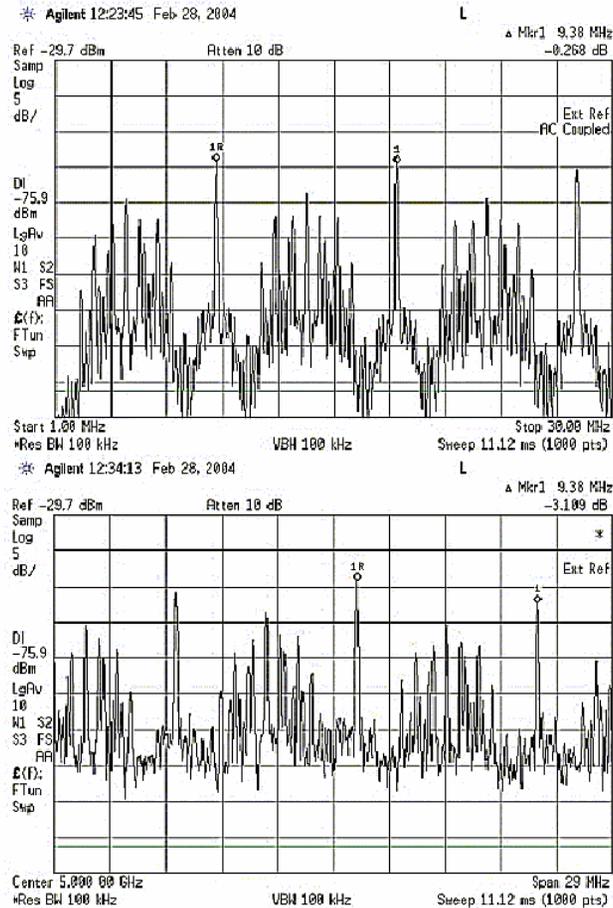


Figure 2: High and low frequency gold spectra with a span 29 MHz and a resolution bandwidth of 100 kHz. The generic features of the spectrum do not change between baseband and 5 GHz.

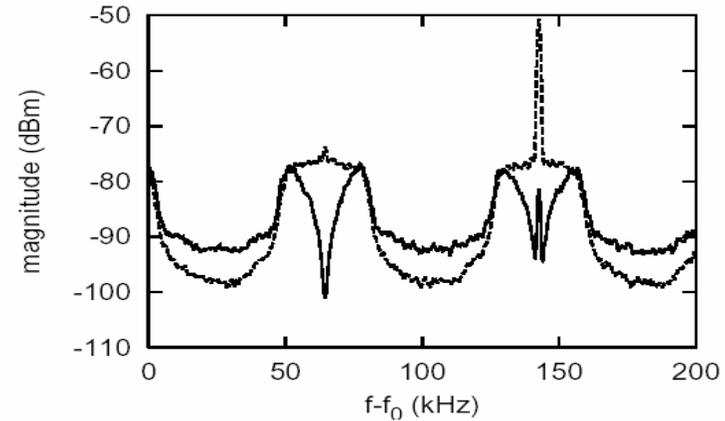
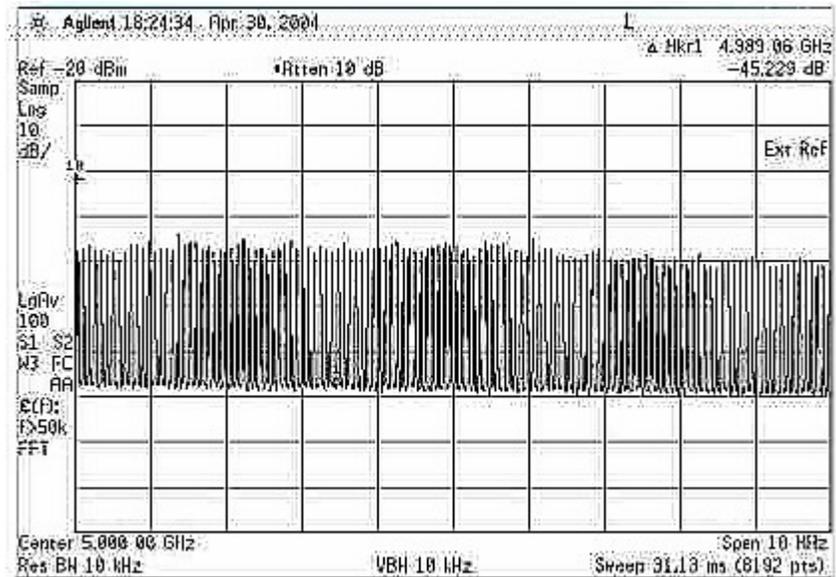


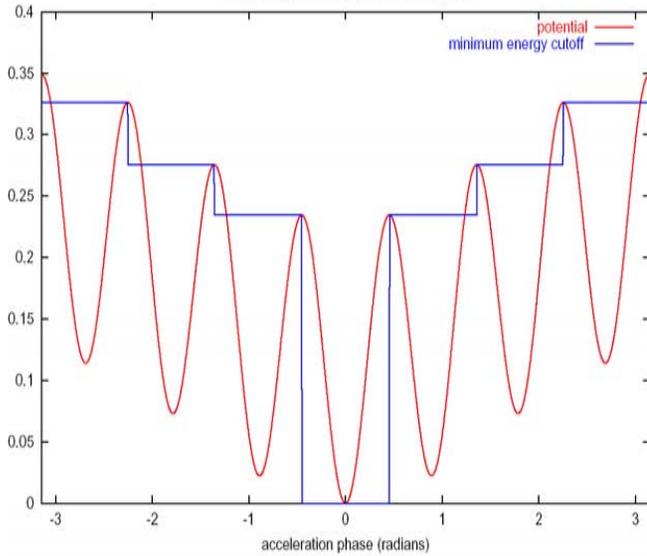
Figure 1: Gold Schottky spectra with (solid) and without (dashed) the one turn delay notch filter,  $f_0 = 4.77$  GHz.



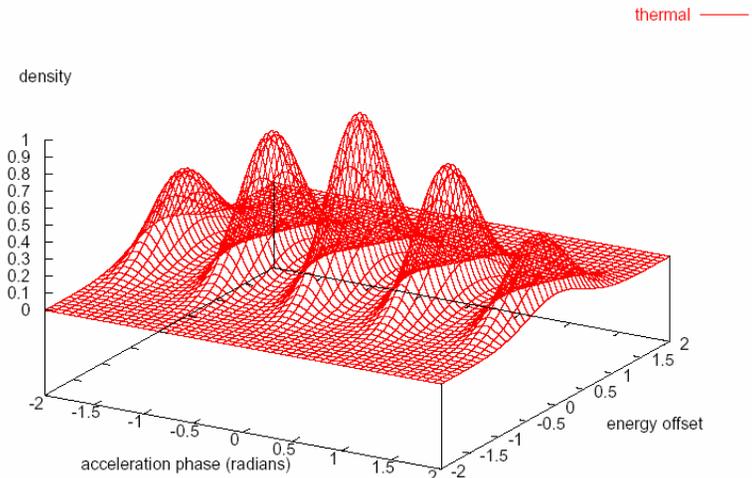
## High frequency proton spectra

# Coherent Line Problem with Gold

4 MV on storage, 0.3 MV on accelerating

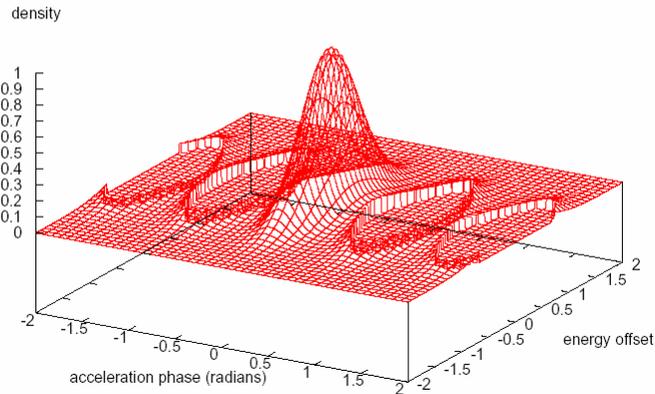


4 MV on storage, 0.3 MV on accelerating

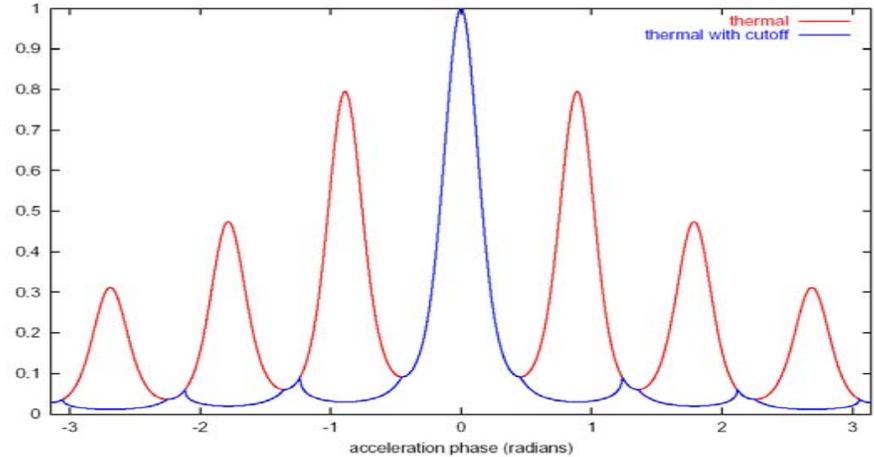


4 MV on storage, 0.3 MV on accelerating

with cutoff



4 MV on storage, 0.3 MV on accelerating



# Coherent Line Problem with Protons

Long lived coherent oscillations produce strong, narrow band spectra.

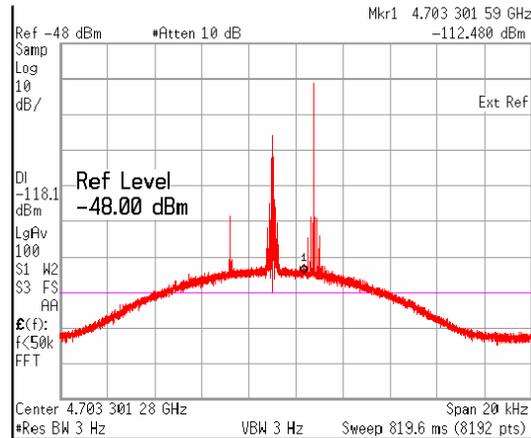
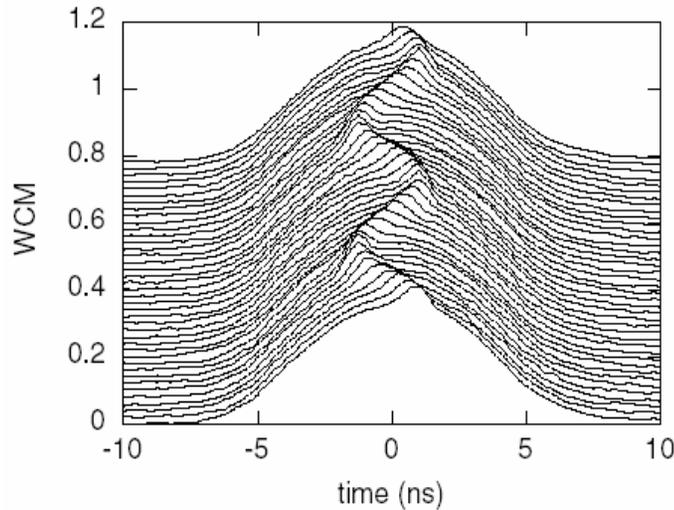


Figure 3: High resolution proton Schottky spectrum.

PHYSICAL REVIEW SPECIAL TOPICS - ACCELERATORS AND BEAMS, VOLUME 7, 044402 (2004)

$$H(\phi, p, s) = p^2/2 + 1 - \cos\phi + \ell\rho(\phi, s). \quad (4)$$

$$p = \sqrt{2J} \cos(\Psi + s), \quad \phi = \sqrt{2J} \sin(\Psi + s).$$

$$-\frac{2a^3\lambda}{\ell} = G_b(r)/r^2 - G_a(r) \approx -0.77r\ln(r)/(1+r), \quad (32)$$

If we take a simple diffusion model, the rate at which a soliton dissipates will scale as  $\nu_{\text{coll}}/a^2$ , with  $a$  being a typical dimension of the soliton. With this extra factor, the dissipation rate for solitons in a gold bunch is about 50 times faster than the dissipation rate for a comparable soliton in a proton bunch.

# Voltage considerations

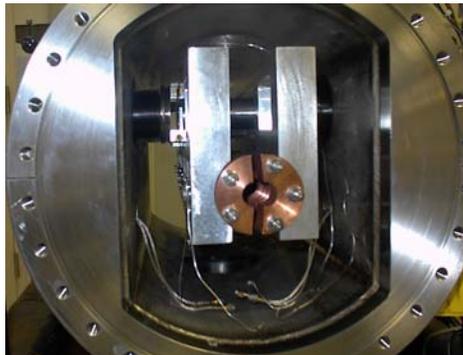
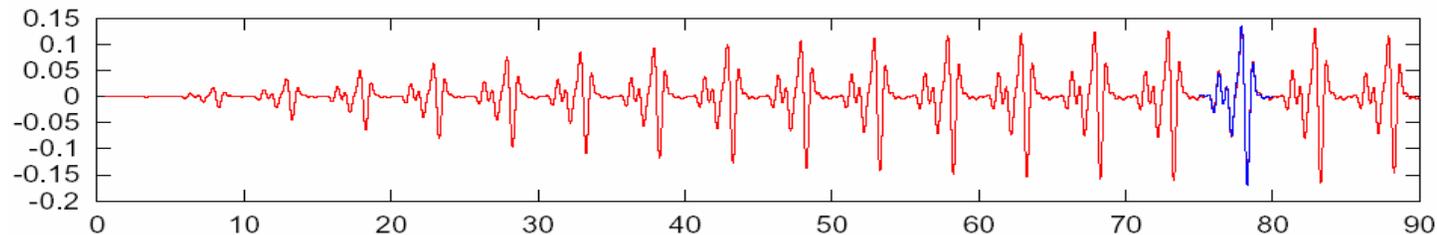
For 4-8 GHz need 3.6 kV rms, large by stochastic cooling standards

Bandwidth-Voltage product sets the cost scale

Bunches are  $\tau_b = 5$  ns long spaced by at least  $T_b = 100$  ns

The value of the kicker voltage matters only when the bunch is present

$$V(t) = \sum_n A_n \sin(2\pi n t / \tau_b + \theta_n) \quad A_n \text{ and } \theta_n \text{ vary smoothly}$$



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## Voltage and Power part 2

Take 21 cavities, 4-8 GHz bandwidth 40 Watts/cavity (10 K each)

$R/Q=100\Omega$ , 10 MHz FWHP bandwidth,  $R \geq 40$  kilo-Ohm

gives 1 to 1.4 kV rms per cavity, or 5.6 kV total

Cavity drive signal needs to be roughly sinusoidal for R (not R/Q) to matter

Suppose  $S_0(t)$  is the drive signal for a broad band kicker (like a resistor).

Periodically extend 
$$S(t) = \sum_{k=0}^N S_0(t - k\tau_b)$$

This creates a signal with 10 MHz  $(1/T_b)$  wide peaks,

spaced by 200 MHz  $(1/\tau_b)$ .

Split and pass through 100 MHz filters, centered on cavity resonance, before power amps. In this way each amplifier sees a piecewise sinusoidal input.

Plan to use fiber optic technology for the delay line filter.

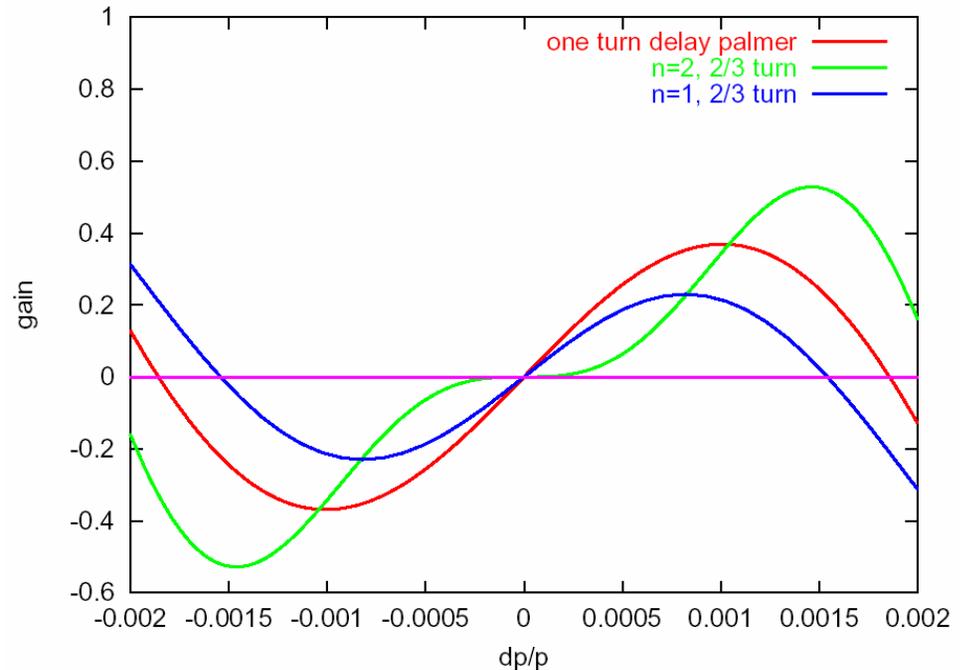
# Low Level Drive for Halo Cooling

For cooling we need a force in reverse to the energy error.

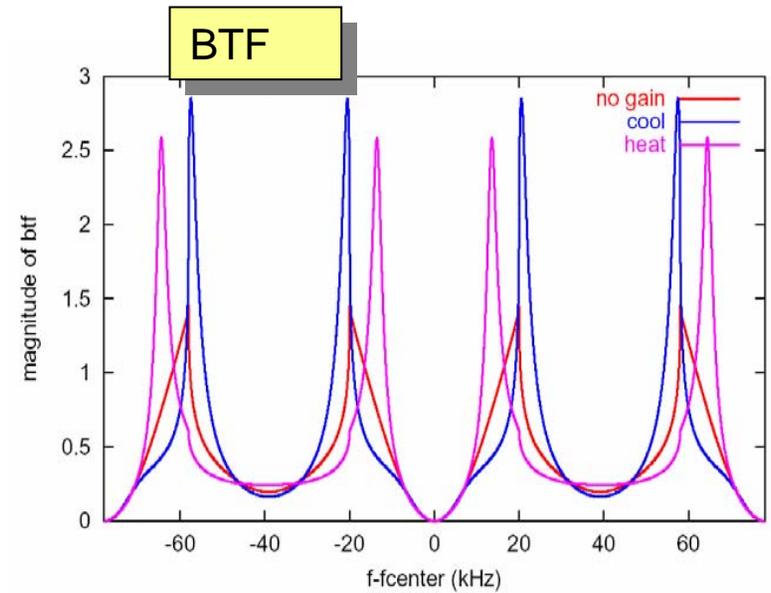
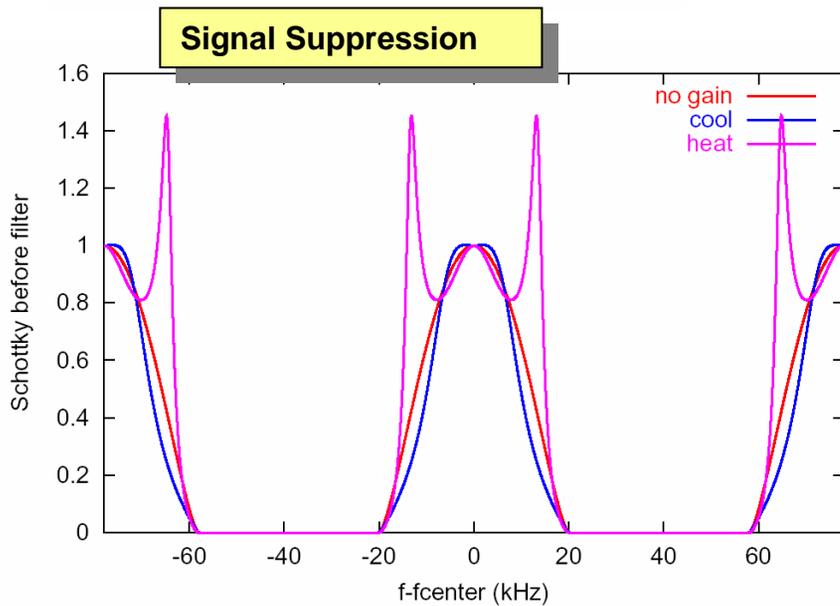
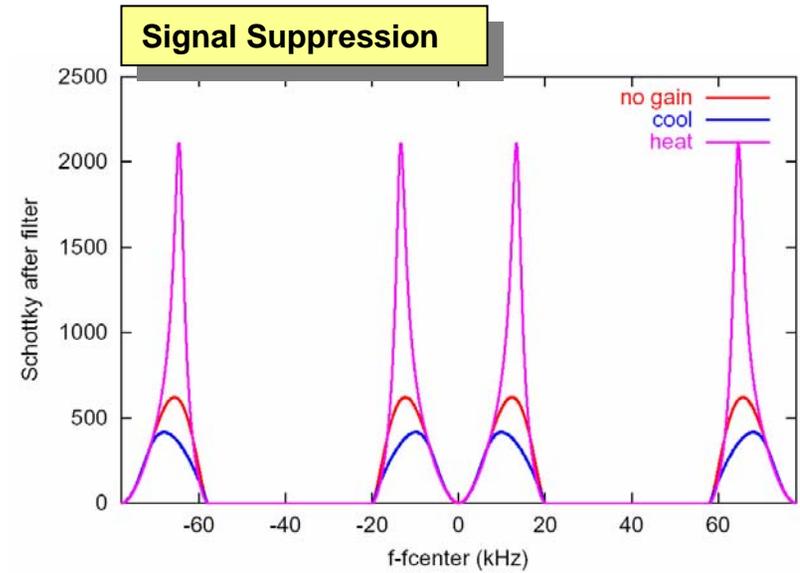
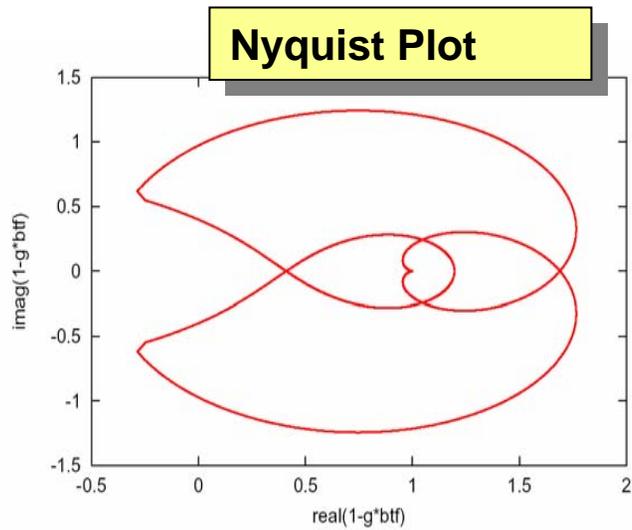
For 4-8 GHz we can use filter cooling

$$S(t) = G(1 - e^{-j\omega T_{rev}})^n I_b e^{j\omega(t - T_d)}$$

For  $n=2$  we cool the full halo.



# System Stability, n=2



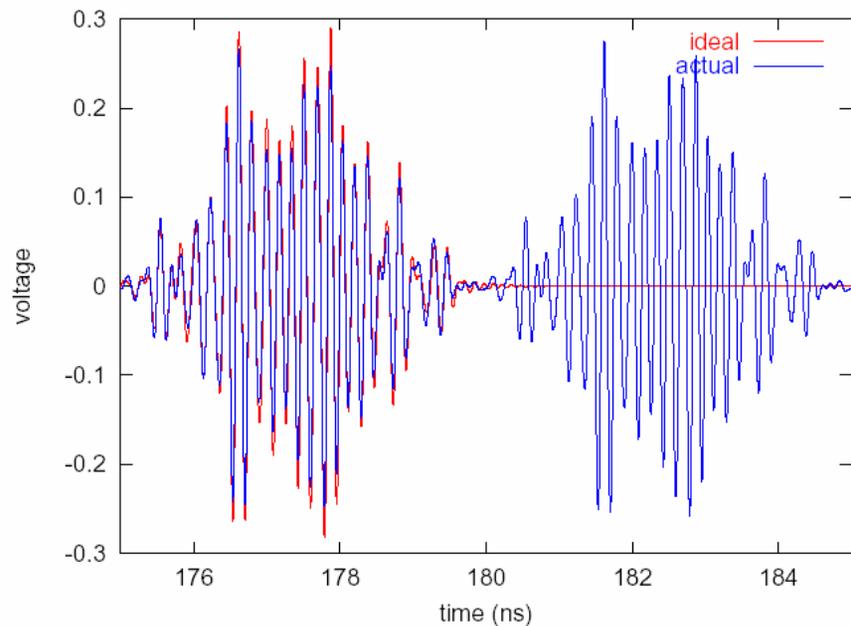
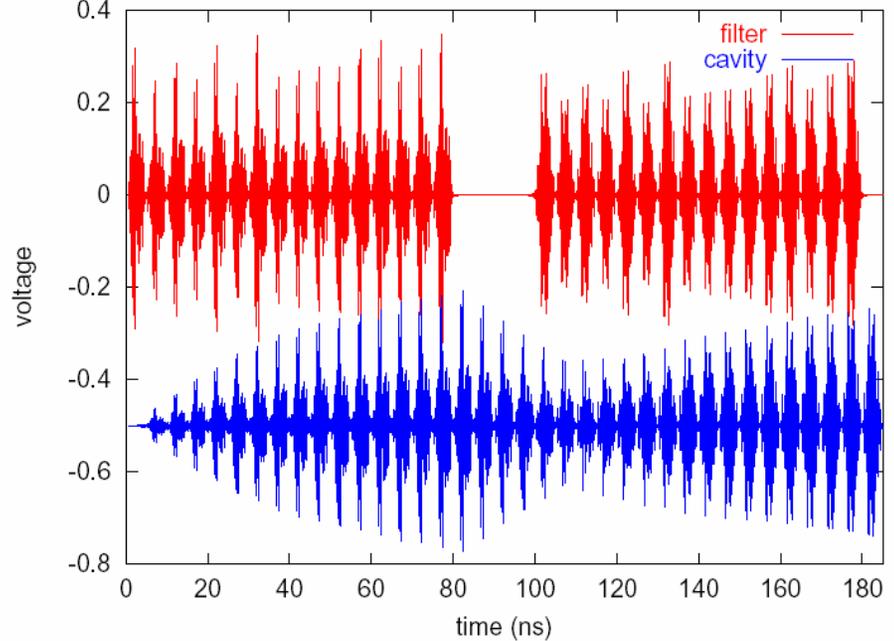
# Error Limit Simulations

Took conservative errors.

- 2 ps timing error
- 20% amplitude errors
- 2 MHz cavity frequency errors

Desired cooling voltage is modeled as band limited noise.

System is well behaved with these errors.



# Plan for this run

- Installing 7-8 GHz and 4-5 GHz kickers in yellow 4 o'clock
- Build 16 channel delay line and full turn delays.
- Have 5 amplifiers so far
- For full comparison with Au beams need  $1.e9$  ions in a witness bunch
- We should get a clean cooling signal albeit at about  $\frac{1}{2}$  the expected rate for the full system.
- Required voltage/cavity very similar to that needed for Au.
- All this assumes rebucketed beam. An experiment without rebucketed beam would be worthwhile.

# Future Plans and Ideas

- Install full 4-8 GHz systems in both rings
- Study interaction with electron cooling
- Consider a wider bandwidth.
  - Cut a chord
  - Use a transverse pickup (dispersion).
  - New kicker technology needed.
- Transverse Cooling

